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~~Transport Equation: Moving Energy on the Photon Train~~ **Heat Transfer - Conduction - Burning Balloons Heat Transfer L1 p4 - Conduction Rate Equation - Fourier's Law** *Heat Transfer: Conduction, Convection, and Radiation* **Heat Transfer L3 p1 - Surface Energy Balance**

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Solution The total energy emitted from the sun $Q = 4\pi R_s^2 \sigma T_s^4$ sun goes equally into all directions, so that $q_{SC}(R) = \frac{Q}{4\pi R^2} = \sigma T_s^4 \left(\frac{R_s}{R}\right)^2$
 $q_{SCmars} = \sigma T_s^4 \left(\frac{R_s}{R_{Mars}}\right)^2$, $q_{SCearth} = \sigma T_s^4 \left(\frac{R_s}{R_{Earth}}\right)^2$
 $q_{SCmars} = 0.4305 \times 1367 = 588 \text{ W/m}^2$

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